## Solving the FMO

Here is the Hamiltonian describing a single excitation in the FMO complex from the latest work by Fleming: A. Ishizaki and G. R. Fleming, Proc. Natl. Acad. Sci. U.S.A. 106, 17255 (2009).

## I. THE FMO HAMILTONIAN FROM PROC. NATL. ACAD. SCI. U.S.A. 106, 17255 (2009)

The units are wavenumbers, i.e. $\mathrm{cm}^{-1}$.

$$
\left.\begin{array}{l}
V=\left(\begin{array}{ccccccc}
12410 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 12530 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 12210 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 12320 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 12480 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 12630 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 12440
\end{array}\right) \\
K
\end{array} \begin{array}{c}
\text { 0 } \\
0
\end{array} \begin{array}{ccccccc}
07.7 & -5.5 & 5.9 & -6.7 & 13.7 & 9.9  \tag{1}\\
87.7 & 0 & -30.8 & -8.2 & -0.7 & -11.8 & -4.3 \\
-5.5 & -30.8 & 0 & 53.5 & 2.2 & 9.6 & -6.0 \\
5.9 & -8.2 & 53.5 & 0 & 70.7 & 17.0 & 63.3 \\
-6.7 & -0.7 & 2.2 & 70.7 & 0 & -81.1 & 1.3 \\
13.7 & -11.8 & 9.6 & 17.0 & -81.1 & 0 & -39.7 \\
9.9 & -4.3 & -6.0 & 63.3 & 1.3 & -39.7 & 0
\end{array}\right)
$$

$$
X=\left(\begin{array}{ccccccc}
-0.2713781 & 0.4351370 & 0.1155647 & -0.0691068 & -0.1655553 & 0.8267552 & 0.0891404 \\
0.2052872 & 0.6427755 & -0.5419359 & 0.4112510 & 0.2403535 & -0.0998430 & -0.1188969 \\
-0.6613888 & -0.2678740 & -0.0013524 & 0.6699533 & 0.0359892 & 0.0090086 & -0.2014764 \\
0.3611684 & -0.3199886 & 0.0562502 & 0.3630684 & 0.4654431 & 0.3438680 & 0.5452517 \\
0.2988589 & 0.1865958 & 0.7071748 & 0.2002269 & 0.2491240 & 0.0240024 & -0.5225318 \\
0.4316540 & -0.3916357 & -0.3371893 & 0.1443333 & -0.4043355 & 0.3766622 & -0.4694742 \\
-0.2017213 & -0.1874828 & -0.2756656 & -0.4295554 & 0.6865415 & 0.2138464 & -0.3828575
\end{array}\right)
$$

$$
\operatorname{diag}(\lambda)=\left(\begin{array}{lllllll}
-143.85474 & -92.22032 & -42.55453 & 0.95012 & 54.82277 & 94.65536 & 128.20133
\end{array}\right)
$$

## A. units

In spectroscopy, the wavenumber $\tilde{\nu}$ of electromagnetic radiation is defined as

$$
\begin{equation*}
\tilde{\nu}=1 / \lambda \tag{2}
\end{equation*}
$$

where $\lambda$ is the wavelength of the radiation in a vacuum. The wavenumber has dimensions of inverse length and SI units of reciprocal meters $\left(m^{-1}\right)$. A wavenumber can be converted into quantum-mechanical energy E in J or regular frequency $\nu$ in Hz according to

$$
\begin{gather*}
E=h c \tilde{\nu}=1.9865 \times 10^{-23} \mathrm{~J} \mathrm{~cm} \times \tilde{\nu}=1.2398 \times 10^{-4} \mathrm{eV} \mathrm{~cm} \times \tilde{\nu},  \tag{3}\\
\nu=c \tilde{\nu}=2.9978 \times 10^{10} \mathrm{~Hz} \mathrm{~cm} \times \tilde{\nu} \tag{4}
\end{gather*}
$$

Note that here wavenumber and the speed of light are in cgs units.

## B. plot

We are interested in the population of the localized sites, given some starting state.

$$
\begin{equation*}
U(t)=X e^{-i t \lambda} X^{\dagger} \tag{5}
\end{equation*}
$$

Let $\psi_{q}$ be the localized wave function for the $q^{\text {th }} \mathrm{BChl}$ in the particle basis and let

$$
\begin{equation*}
\Pi_{i} \psi_{q}=\delta_{i q} \tag{6}
\end{equation*}
$$

We start the system in particle basis $\psi_{i}$ and want the probability of finding the particle in the $q^{\text {th }} \mathrm{BChl}$

$$
\begin{equation*}
\operatorname{Prob}_{q}(t)=\operatorname{Tr}\left(\Pi_{q}\left|U(t) \psi_{i}\right\rangle\left\langle\psi_{i} U(t)\right|\right)=\operatorname{Tr}\left(\Pi_{q} U(t) \Pi_{i} U^{\dagger}(t)\right)=\left|\left\langle\psi_{q} \mid U(t) \psi_{i}\right\rangle\right|^{2} \tag{7}
\end{equation*}
$$

Time evolution of the population at each BChl of the FMO complex


FIG. 1: Time evolution with initial state BChl-1 |1>.

